Electromagnetic Waves in a Uniform Gravitational Field and Planck’s Postulate

Luis Acedo* and Michael M. Tung
Instituto de Matemática Multidisciplinar,
Universitat Politècnica de València,
Camino de Vera s/n, 46022 Valencia, Spain

May 27, 2015

Abstract

The gravitational redshift forms the central part of the majority of the classical tests for the general theory of relativity. It could be successfully checked even in laboratory experiments on the earth’s surface. The standard derivation of this effect is based on the distortion of the local structure of spacetime induced by large masses. The resulting gravitational time-dilation near these masses gives rise to a frequency change of any periodic process, including electromagnetic oscillations as the wave propagates across the gravitational field. This phenomenon can be tackled with classical electrodynamics assuming a curved spacetime background and Maxwell’s equations in a generally covariant form. In the present paper, we show that in a classical field-theoretical context the gravitational redshift can be interpreted as the propagation of electromagnetic waves in a medium with corresponding conductivity \( \sigma = g / (\mu_0 c^3) \), where \( g \) is the gravitational acceleration and \( \mu_0 \) is the vacuum magnetic permeability. Moreover, the energy density of the wave remains proportional to its frequency in agreement with Planck’s postulate.

*Corresponding author e-mail: luiacrod@imm.upv.es
1 Introduction

The first classical test of the general theory of relativity in a terrestrial environment was devised and carried out by Robert Pound and Glen A. Rebka, Jr., in 1959 [1]. The idea of this experiment is quite simple: An atom of iron-57 emits by radioactive decay a gamma photon with an energy of 14 keV. Such a photon travels upward (in this case from the basement of Jefferson Laboratory at Harvard’s University) until it reaches a certain height $h$ (in the original experiment at the attic of the building it was $h = 22.5$ m) where its energy has changed to

$$E_1 = E_0 \left(1 - \frac{gh}{c^2}\right),$$

where $g$ denotes the strength of the gravitational field, being equivalent to the acceleration of objects under its influence. The interpretation of the redshift in a static gravitational field is not as simple as it appears; for a detailed discussion see Ref. [2].

At the maximum height of its trajectory, the redshifted photon can no more be absorbed by the iron-57 atoms in the receiver. In a moment of inspiration, Pound and Rebka realised that the gravitational redshift of the photon can be cancelled out by an artificially created Doppler blueshift equivalent to a downward motion towards the basement:

$$E_2 = E_1 \left(1 + \frac{v}{c}\right) = E_0 \left(1 + \frac{v}{c} - \frac{gh}{c^2} + O\left(c^{-3}\right)\right).$$

As can be clearly seen, the cancellation with $E_2 \approx E_0$ takes place if we choose the velocity $v = gh/c$. However, the fractional energy change $(E_1 - E_0)/E_0 = 2.5 \times 10^{-15}$ is so small that the experiment would be contaminated by atom recoiling. Fortunately, just a year before, Rudolf L. Mössbauer had discovered that atoms belonging to a solid lattice share the recoiling momentum and, consequently, the associated energy and velocity change is negligible [3]. By using Mössbauer spectroscopy, Pound and Rebka were then able to measure the velocity necessary to counteract the gravitational redshift by the Doppler effect, $v = gh/c \approx 7.5 \times 10^{-4}$ mm/s. These very small velocities were obtained by placing the sample on a conical speaker membrane reproducing a low-frequency sound between 10 and 50 Hz. Note that as a sequel to this initial experiment, many higher-precision tests were added, e.g. in 1980 by means of a space-borne hydrogen maser [4].
From a historical point of view, already in 1784, John Michell, an English philosopher and geologist, anticipated the gravitational weakening of starlight in a letter to Henry Cavendish [5]. Michell’s calculations were based on the Newtonian corpuscular theory of light which was later on rejected after the arrival of the wave theory in the early XIXth century.

The objective of this paper is to analyse the propagation of light in a gravitational field by means of the covariant Maxwell equations with a curved spacetime background. We will consider a uniform gravitational field and a uniform accelerating frame as proposed by Desloge [6] and assume the following line element:

$$ds^2 = -\alpha^2(z) c^2 dt^2 + dx^2 + dy^2 + dz^2,$$

where time-dilation is included either by considering a uniform gravitational field or by a uniform accelerating frame along the \(z\) axis:

$$\alpha(z) = \begin{cases} e^{gz/c^2} & \text{uniform gravitational field (UGF)} \\ 1 + \frac{gz}{c^2} & \text{uniform accelerating rigid frame (UAF)} \end{cases}$$

The UGF metric is a straightforward solution of the condition that the initial local acceleration of a particle must have the same value at all points in order to describe a uniform field. The corresponding differential equation therefore is:

$$\frac{1}{\alpha} \frac{d\alpha}{dz} = \frac{g}{c^2}, \quad \text{with} \quad \alpha(0) = 1.$$  

On the other hand, the UAF metric for a uniformly accelerating rigid frame in field-free space assumes an underlying flat space, which yields the following simple differential equation and boundary conditions:

$$\frac{d^2 \alpha}{dz^2} = 0, \quad \frac{d\alpha}{dz}(0) = \frac{g}{c^2}, \quad \alpha(0) = 1.$$  

It is not difficult to see that the particular expressions for \(\alpha(z)\) in Eq. (4) are just the exact solutions of these two differential systems, corresponding to the UGF and UAF case, respectively.

The UAF metric describes an underlying flat spacetime, whereas the UGF metric represents curved spacetime as expected. So in principle, the two approaches of Eq. (4) are fundamentally different and were used by Desloge to explicitly show that the observations made in a UGF are not strictly identical to those of a UAF. Here, the statement of the principle of equivalence between acceleration...
and gravity is only valid as a heuristic approximation. In this context, the two different metrics also allowed to analyse and discuss gravitational redshift effects in a straightforward manner. 

In our opinion, it would be pedagogical and helpful to gain new insights if this approach were also extended to include the effects of gravitational redshift from the point of view of classical electrodynamics in combination with general relativity. Classical references on the subject arrive at the covariant formulation of Maxwell’s equations and stop there. Only in some of the more advanced textbooks particular solutions are fully discussed. We will show that Desloge’s metric is an excellent testing ground by obtaining explicit solutions of Maxwell equations for electromagnetic waves in a curved spacetime background. Furthermore, we will see how the conceptual transition from classical electrodynamics to its extension in general relativity is minimised in this approach.

The paper is organised as follows: In Section 2 we first set up the Maxwell equations for the UGF and UAF metrics. In particular, we study the electrostatic field of an infinite and uniformly charged plate in a uniform gravitational field and derive its solution. Next, in Section 3, the electromagnetic wave equation is formulated for a general UGF frame. An approximate analytical solution of the wave equation is also calculated and expressed in terms of the two linearly independent solutions $A_i(z)$ and $B_i(z)$ of the Airy equation $y''(x) - xy = 0$ extended to the complex plane by analytic continuation. It can be shown that the relation among the energy density of the wave and its frequency satisfies Planck’s postulate. Section 4 concludes the paper with some final remarks and observations.

2 Covariant Maxwell Equations in a Uniform Gravitational Field

In 4-dimensional spacetime, the Faraday tensor, or covariant electromagnetic field tensor, allows the physical laws which govern electromagnetic phenomena to be written in a very concise form. For an underlying metric with signature $(−, +, +, +)$, it is defined by

$$F_{\mu\nu} = \begin{pmatrix}
0 & -E_x/c & -E_y/c & -E_z/c \\
E_x/c & 0 & B_z & -B_y \\
E_y/c & -B_z & 0 & B_x \\
E_z/c & B_y & -B_x & 0
\end{pmatrix}, \quad (5)$$
where as usual the electromagnetic field is decomposed into the field vectors $\mathbf{E} = (E_x, E_y, E_z)$ and $\mathbf{B} = (B_x, B_y, B_z)$ as seen in a frame of a particular observer. Using Desloge’s approach, the inverse of $F^{\mu\nu}$ with the metric of Eq. (3) is given in contravariant form by

$$F^{\mu\nu} = \begin{pmatrix}
0 & E_x/c\alpha^2(z) & E_y/c\alpha^2(z) & E_z/c\alpha^2(z) \\
-E_x/c\alpha^2(z) & 0 & B_z & -B_y \\
-E_y/c\alpha^2(z) & -B_z & 0 & B_x \\
-E_z/c\alpha^2(z) & B_y & -B_x & 0
\end{pmatrix}. \tag{6}$$

Maxwell’s equations can then be recast in covariant form. For this purpose the source equations of the electric and magnetic fields are summarised in a single relation containing the covariant derivative of the electromagnetic field tensor:

$$F^{\mu\nu},_{\nu} = \mu_0 j^\mu, \tag{7}$$

where the semicolon denotes the covariant derivative and $j^\mu = (\rho, j_x, j_y, j_z)$ is the current four-vector. The rotational equation for the electric field and the divergence-free condition of the magnetic field are incorporated in the cyclic equation for $F^{\mu\nu}$:

$$F_{[\lambda\mu;\nu]} = 0 \quad \text{or} \quad F_{\lambda\mu;\nu} + F_{\mu\nu;\lambda} + F_{\nu\lambda;\mu} = 0, \tag{8}$$

where the cyclic permutations of the indices may be abbreviated by the common bracket notation. This implies that in all terms the covariant derivatives, which include Christoffel symbols, cancel out, and we can replace the covariant derivatives by ordinary derivatives obtaining

$$F_{[\lambda\mu;\nu]} = 0 \quad \text{or} \quad F_{\lambda\mu;\nu} + F_{\mu\nu;\lambda} + F_{\nu\lambda;\mu} = 0, \tag{9}$$

where the comma denotes now conventional partial derivatives. After expanding the covariant derivative in Eq. (7), we also find

$$F^{\mu\nu},_{\nu} + \frac{\sqrt{-g}}{\sqrt{-g}} F^{\mu\nu} = \mu_0 j^\mu, \tag{10}$$

where we have taken into account that the contraction of the symmetric Christoffel symbols with the antisymmetric electromagnetic tensor is zero in the absence of torsion, namely $\Gamma^\mu_{\alpha\beta} F^{\alpha\beta} = 0$. In Desloge’s approach, the metric tensor is given
by

\[
g_{\mu\nu} = \begin{pmatrix}
-\alpha^2(z) & 0 & 0 & 0 \\
0 & 1 & 0 & 0 \\
0 & 0 & 1 & 0 \\
0 & 0 & 0 & 1 \\
\end{pmatrix},
\]

and consequently, \( \det (g_{\mu\nu}) = -\alpha^2(z) \). From Eqs. (6) and (10) we arrive after some simplification at the following conditions for the electric field

\[
\nabla \cdot E - \frac{\dot{\alpha}}{\alpha} E_z = \frac{\rho}{\varepsilon_0} \alpha^2(z),
\]

and similarly for the magnetic field

\[
\nabla \times B = \mu_0 j + \frac{\dot{\alpha}}{\alpha} B \times \hat{k} + \frac{1}{c^2 \alpha^2} \frac{\partial E}{\partial t},
\]

where \( \hat{k} \), as usual, denotes the unit vector in \( z \)-direction. In an analogous way, from Eq. (8) we may derive the two remaining Maxwell equations:

\[
\nabla \times E = -\frac{\partial B}{\partial t},
\]

\[
\nabla \cdot B = 0.
\]

Before starting to find the full electromagnetic wave solution of this system, it is useful to study the simpler electrostatic case. For this purpose, consider in a particular UGF frame an infinite metallic plate which is uniformly charged and is perpendicular to the \( z \)-axis. From Eqs. (12) and (14) we find that the electric field is irrotational and satisfies

\[
\nabla \cdot E - \frac{g}{c^2} E_z = 0, \quad \text{(outside the charged plate)}
\]

which indicates that the translational Poincaré symmetry has been broken by gravitation. As the electric field is aligned in the \( z \)-direction, in the vacuum Eqs. (12) and (16) reduce to

\[
\frac{\partial E_z}{\partial z} - \frac{g}{c^2} E_z = 0, \quad \text{(UGF)}
\]

\[
\frac{\partial E_z}{\partial z} - \frac{g/c^2}{1 + g z/c^2} E_z = 0, \quad \text{(UAF)}
\]
which can be readily integrated to yield

\[ E_z(z) = E_0 e^{z/L}, \quad \text{(UGF)} \]  
\[ E_z(z) = E_0 \left( 1 + \frac{z}{L} \right). \quad \text{(UAF)} \]

Here \( L = c^2/g \) is a characteristic length scale associated with the gravitational field. For weak gravitational fields this scale is very large: Assuming, for example, a uniform gravitational field with an acceleration corresponding to the local acceleration at the surface of the earth, namely \( g = 9.8 \text{ m/s}^2 \), produces the scale value \( L \approx 0.97 \text{lyr} \). The length scale in general also provides for an estimate in which domain the UGF and UAF descriptions agree or differ in their predictions. In the domain \( z \ll L \) expression Eq. (19) converges to Eq. (20), so that both results agree to first order:

\[ E_z(z) = E_0 \left[ 1 + \frac{z}{L} + O\left(\left(\frac{z}{L}\right)^2\right) \right]. \] (21)

As it has been shown for other phenomena, a uniformly accelerated rigid frame and a uniform gravitational field are not strictly equivalent on larger scales, although the equivalence principle, which guided Einstein heuristically towards a formulation of general relativity, is still valid locally for weak fields and small accelerations [6, 7].

It is worthwhile to note that the gravitational acceleration corresponding to Eqs. (19) and (20) are not exactly realised in nature. In fact, the discrepancy between the two alternatives (UGF or UAF) occurs outside the physical domain. The physical domain is fixed by the scale \( L = c^2/g \) such that \( |z| \ll L \), since the space-time metric of the UAF description strictly applies only in the limit \( g z/c^2 \ll 1 \), and furthermore the UGF approach becomes problematic in the sufficiently large- \( z \) domain. In any case, the full analytical treatment of the considerably more complicated case with a black-hole background spacetime, which represents a physical and very strong gravitational field in its vicinity, deserves special attention and is planned in a future work. Nevertheless, the current approach serves as a viable and instructive guide to explore wave-like solutions of Maxwell’s solutions and their energy content for uniformly accelerating frames.
3 Electromagnetic waves in a uniform gravitational field

Considering Desloge’s UGF metric of Eq. (4), which implies a uniform gravitational acceleration along the $z$-axis, Maxwell equations in vacuum take the following form:

$$\nabla \cdot \mathbf{E} = \frac{g}{c^2} E_z, \quad (22a)$$

$$\nabla \times \mathbf{E} = \frac{\partial \mathbf{B}}{\partial t}, \quad (22b)$$

$$\nabla \cdot \mathbf{B} = 0, \quad (22c)$$

$$\nabla \times \mathbf{B} = \frac{g}{c^2} \mathbf{B} \times \mathbf{k} + \frac{1}{c^2} e^{-2gz/c^2} \frac{\partial \mathbf{E}}{\partial t}. \quad (22d)$$

In order to derive the wave equation, we apply the standard technique in classical electrodynamics by taking the curl of Eq. (22b) and thereby obtain

$$\nabla \times (\nabla \times \mathbf{E}) = \nabla (\nabla \cdot \mathbf{E}) - \nabla^2 \mathbf{E} = -\frac{\partial}{\partial t} \nabla \times \mathbf{B}. \quad (23)$$

By direct substitution of the expressions for the divergence of the electric field, Eq. (22a), and the curl of the magnetic field, Eq. (22d), we finally get

$$\nabla^2 \mathbf{E} - \frac{g}{c^2} \frac{\partial}{\partial t} \mathbf{B} \times \mathbf{k} - \frac{1}{c^2} e^{-2gz/c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} - \frac{g}{c^2} \nabla E_z = 0. \quad (24)$$

Proceeding in a similar way, we take the curl of the curl of the magnetic field and simplify by using the remaining Maxwell equations to arrive at

$$\nabla^2 \mathbf{B} + \frac{g}{c^2} \frac{\partial \mathbf{B}}{\partial z} + \frac{2g}{c^4} \frac{\partial}{\partial t} \mathbf{E} \times \mathbf{k} - \frac{1}{c^2} e^{-2gz/c^2} \frac{\partial^2 \mathbf{B}}{\partial t^2} = 0. \quad (25)$$

Equations (23) and (25) are apparently quite different from the well-known wave equations. In order to explain the behaviour of electromagnetic waves in the UGF system, we make some simplifying assumptions:

(i) The wave travels upwards or downwards, parallel to the $z$-axis.

(ii) The are no longitudinal electromagnetic components: $E_z = 0, B_z = 0.$
(iii) The electric, magnetic and propagation vector satisfy the standard right-hand-rule which implies \( E = cB \times \hat{k} \) and \( B = -E \times \hat{k}/c \).

With these conditions, Eq. (24) becomes

\[
\nabla^2 E - \frac{g}{c^3} \frac{\partial E}{\partial t} - \frac{1}{c^2} e^{-2qz/c^2} \frac{\partial^2 E}{\partial t^2} = 0.
\]

(26)

Except for the additional factor of the second-order time derivative, this equation coincides with the telegraph equation for the propagation of electromagnetic waves in a conducting medium [8]. As in the telegraph equation, we identify in Eq. (26) the coefficient of the first time derivative with \( \mu_0 \sigma \), so that the conductivity of the gravitational field can be taken as \( \sigma = g/(\mu_0 c^3) \). We now propose the following general solution for the complex electric field:

\[
E = E_0 e^{\gamma(z) - i\omega t},
\]

(27)

where \( i \) is the imaginary unit, \( \omega \) is the frequency of the wave, \( t \) is the coordinate time, and \( \gamma(z) \) is a function still to be determined. Observe that, if we use coordinate time instead of local time at a fixed spatial position, the frequency measured is constant. By inserting Eq. (27) into Eq. (26) we obtain

\[
\frac{d^2 \gamma}{dz^2} + \left( \frac{d\gamma}{dz} \right)^2 + \frac{\omega^2}{c^2} e^{-2qz} + \frac{i q \omega}{c} = 0,
\]

(28)

where \( q = g/c^2 \) is the inverse of the characteristic length of the gravitational field. Equation (26) is a second-order non-linear differential equation which, fortunately, can be linearised by the variable change \( \gamma(z) = \ln (z + A(z)) \), with the new unknown function \( A(z) \). It then follows that

\[
\frac{d^2 A}{dz^2} + \left( \frac{\omega^2}{c^2} e^{-2qz} + \frac{i q \omega}{c} \right) (A + z) = 0.
\]

(29)

This differential equation for \( A(z) \) is non-homogeneous, however one of its particular solutions is simply \( A_p(z) = -z \). The general solution of the homogeneous equation, \( A_h(z) \), can be expressed in terms of Bessel functions with complex index and the gamma function with complex argument [9, p. 447]:

\[
A_h(z) = \kappa_1 J_{\nu e^{3\pi i/4}} \left( \nu^2 e^{-z_i/\nu^2} \right) \Gamma \left( 1 + e^{3\pi i/4} \nu \right)
+ \kappa_2 J_{-\nu e^{3\pi i/4}} \left( \nu^2 e^{-z_i/\nu^2} \right) \Gamma \left( 1 - e^{3\pi i/4} \nu \right).
\]

(30)
Figure 1: Electric field amplitude as a function of scaled height, $z_\star$, for an electromagnetic wave in a strong uniform gravitational field with $\nu = 15$. The coefficients of $A_h(z)$ are $\kappa_1 = 10^{13}$ and $\kappa_2 = 0$.

Here we have abbreviated $\nu = \sqrt{\omega/qc}$ and $z_\star = \omega z/c$. The parameter $\nu$ is usually very large for typical frequencies and gravitational accelerations. The electric field of the wave is then given by

$$E(z, t) = E_0 \left( A_h(z) + A_p(z) + z \right) e^{-i\omega t} = E_0 A_h(z) e^{-i\omega t},$$

(31)

because the particular solution of the non-homogenous equation, $A_p(z)$ cancels out. In Fig. 1 we have plotted the electric field amplitude, which is obtained from the real part of Eq. (31) after substituting Eq. (30). The numerical parameters are chosen $\nu = 15$, $\kappa_1 = 10^{13}$, and $\kappa_2 = 0$.

The damping of the wave is associated with the gravitational redshift of photons as discussed below. Notice that the wavelength is also increasing and, reciprocally, the frequency is decreasing as the wave travels upwards through the uniform gravitational field. The coefficient $\kappa_2$ must be zero, because the second
term in Eq. (30) corresponds to an amplification of the wave as it travels upwards and, consequently, is unphysical.

On the other hand, as discussed in Sec. 1, gravitational redshifts could be successfully detected even for weak gravitational fields, such as the local field at the surface of the earth. This justifies to also study the approximation of the general solution Eq. (30) for the case of weak fields, with the limits $g \to 0$ or $\nu \to \infty$. If we try to accomplish this task directly from Eq. (30), we will face some technical difficulties because of the imaginary index of the Bessel functions. Moreover, the method of the stationary phase is also difficult to apply because $\nu$ appears also as an argument of the Bessel functions. It is far more convenient to start with the differential equation for $A(z)$, given by Eq. (29), and carry out the expansion for $q = g/c^2 \ll 1$:

$$\frac{d^2 A}{dz^2} + \left(\frac{\omega^2}{c^2} (1 - 2qz) + i\frac{q\omega}{c}\right) (A + z) = 0. \quad (32)$$

An explicit solution of the homogeneous equation, Eq. (32), is now found in terms of the Airy functions

$$A_h(z) = \kappa_1 \text{Ai} \left(\frac{2z_* - \nu^2 - i}{(2\nu)^{2/3}}\right) + \kappa_2 \text{Bi} \left(\frac{2z_* - \nu^2 - i}{(2\nu)^{2/3}}\right), \quad (33)$$

where $z_* = \omega z/c$ and $\nu = \sqrt{\omega c/g}$, as before. Again, out of physical grounds, one has to take $\kappa_2 = 0$. Moreover, it can then be seen that only the imaginary part of the Airy function of the first kind in Eq. (33) is physically meaningful, because it corresponds to the damping of the wave travelling along the $z$-axis in positive direction. The result is shown in Fig. 2 for $\nu = 20$.

A further simplification may be obtained for $\nu \to \infty$ by using the following asymptotic expansion of the Airy function for large $|\zeta|$ with $\zeta \in \mathbb{C}$, see Ref. [9, p. 448, eq. 10.4.59]:

$$\text{Ai}(\zeta) = \frac{1}{2\sqrt{\pi}} \zeta^{-1/4} e^{-\frac{2}{3} \zeta^{3/2}} \sum_{k=0}^{\infty} (-1)^k c_k \left(\frac{2}{3}\zeta^{3/2}\right)^{-k}, \quad (| \arg \zeta | < \pi) \quad (34)$$

where $c_0 = 1$ and $c_k = \Gamma(3k + 1/2)/\left(54^{k} k! \Gamma(k + 1/2)\right)$.

As we are interested in recovering the classical result for gravitational redshifts in a weak gravitational field, we can safely ignore the algebraic prefactors in Eq. (34). The reason for this simplification is as follows: the vertical distance between the emission point of the photon and the receiver in a Pound-Rebka type
Figure 2: Electric field amplitude as a function of scaled height, $z_*$, for an electromagnetic wave in a strong uniform gravitational field with $\nu = 20$. We have used the approximation in Eq. (33) with $\kappa_1 = 1$ and $\kappa_2 = 0$. 
The experiment is much smaller than the characteristic length associated with the approximately uniform field, \( L = c^2/g \gg z \). This implies that \( z_\ast \ll \nu^2 \), and consequently

\[
\zeta = \frac{2z_\ast - \nu^2 - i}{(2\nu)^{2/3}} = \left( \frac{2z_\ast}{\nu^2} - 1 - \frac{i}{\nu^2} \right) \frac{\nu^{4/3}}{2^{2/3}} = \frac{\nu^{4/3}}{2^{2/3}} \left( -1 - \frac{i}{\nu^2} + O(z_\ast/\nu^2) \right).
\]

The terms \( O(z_\ast/\nu^2) \) can then be regarded as the prefactors multiplying the exponential term in Eq. (34).

Ignoring all \( \nu \)-dependent factors which later on can be absorbed into the definition of \( E_0 \) and taking both possible roots in the exponential, Eq. (32) reduces to

\[
A_h(z) \sim e^{\pm \frac{i}{2} \nu^2 [1 + (i - 2z_\ast)/\nu^2]^{3/2}} \approx e^{\pm \frac{i}{2} \nu^2 [1 + \frac{3}{2}(i - 2z_\ast)/\nu + O(\nu^{-6})]^{3/2}},
\]

where we have expanded the exponent as a series of powers of \( \nu^{-2} \) up to second order. Finally, after choosing the negative sign and substituting \( z_\ast/\nu^2 = g z/c^2 \), Eq. (31) gives the following result for a damped wave

\[
E = E_0 e^{-g z/c^2} e^{i \omega (z/c - t)}.
\]

Notice that \( t \) is the coordinate time as introduced in Eq. (27). Here we have again included all remaining terms containing \( \nu \) (but not in combination with \( z_\ast/\nu^2 \)) in the amplitude \( E_0 \). A similar expression may be derived for the magnetic field. This solution represents a damped electromagnetic wave in a conducting medium with conductivity \( \sigma = g/(c^2 \mu_0) \), as could have been anticipated by inspection of Eq. (26). Due to the factor \( e^{-2gz/c^2} \) in the second-order time derivative in Eq. (26), the second-order approximation contains a variable frequency \( \omega \).

Knowing the explicit form of the electric and magnetic fields put us in the position to be able to calculate the energy density of the wave as a temporal average over the coordinate time:

\[
\rho(z) = \frac{1}{2} \left( \varepsilon_0 E^2 + \frac{B^2}{\mu_0} \right) = \frac{1}{2} \left( \varepsilon_0 E_0^2 + \frac{B_0^2}{\mu_0} \right) e^{-gz/c^2} \left\langle \cos^2 \left( \omega (z/c - t) \right) \right\rangle = \rho_0 e^{-gz/c^2}.
\]

Obviously the choice of time variable (coordinate or proper time) for the temporal average in Eq. (37) can not influence the result for the energy density. If \( T \) denotes the wave period by a static clock in the coordinate frame, then the proper wave period in the UGF frame is \( \mathcal{T} = \sqrt{-g_{00}} T \). It is then easy to see that the average
in both frames for one wave period $T$ or $\mathcal{T}$, respectively, yields the same factor $1/2$.

Note that the same exponential factor also appears in the expression for the frequency due to standard gravitational redshift

$$\omega(z) = \frac{\omega_0}{\sqrt{-g_{00}}} = \omega_0 e^{-gz/c^2}, \quad (38)$$

where we have used the metric Eq. (11) in the UGF system defined in Eq. (4).

Therefore, Eqs. (37) and (38) demonstrate that the ratio of energy density and frequency of the wave travelling through the uniform gravitational field is always constant, regardless of its height $z$ as measured by a static receiver at this position. If $n$ is the average number of photons per unit volume, their corresponding energy density is $\rho = n\hbar\omega$ according to Planck’s fundamental postulate of quantum mechanics. It is then clear that

$$\frac{\rho}{\omega} = \frac{\rho_0}{\omega_0} = n\hbar. \quad (39)$$

Hence, we observe that general relativity is compatible with Planck’s postulate concerning the interpretation of the redshift in a strong uniform gravitational field from the point of view of the covariant Maxwell equations in a curved spacetime.

## 4 Conclusions and Remarks

In this paper we have studied the solutions of Maxwell equations in a uniform gravitational background field or, alternatively, in a uniform accelerating rigid frame. We have shown that explicit solutions can be found for electrostatic fields produced by an evenly charged metallic plate and for the case of electromagnetic waves in the vacuum. The wave equation in a gravitational field is analogous to the telegraph equation obtained in classical electrodynamics when electromagnetic waves propagate in a conducting medium. However, the conductivity $\sigma = g/(\mu_0 c^2)$ of such a medium, caused for example by the gravitational field of a typical celestial body, is extremely small. For any feasible test frequency $\omega$ of light, this corresponds to a very large penetration depth of the associated electromagnetic waves.

Moreover, we must recognise that uniform gravitational fields are an idealised case and as such not found in nature. The Schwarzschild metric would be the adequate framework to study realistic gravitational fields. Nevertheless, even as a local approximation for the field near the surface of the earth, the solutions found
provide a useful insight into the behaviour of electromagnetic waves and photons in a gravitational field. In particular, we could show that the amplitude is described in terms of the Airy function depending on height.

The ratio among the energy density of the electromagnetic wave and its frequency is fixed to a constant as the wave travels across the gravitational field. This constant is proportional to Planck’s constant, which is to be expected by compatibility with Planck’s fundamental postulate. Any other result would pose a serious contradiction between the general theory of relativity and quantum mechanics. A similar coherence between both theories is also found for the Compton and Doppler effect in special relativity as studied from a kinematic point of view (taking into account the recoil of a massive body which emits photons) [10, pp. 194].

The relation among quantum phenomena and gravitation in earth-bound experiments goes beyond mere academic or pedagogical interest. For example, the recent proposal for a probabilistic description of gravity, the so-called entropic theory of gravity [11], has been argued to fail for the description of the aforementioned experimental results for quantum states of ultracold neutrons in the earth’s gravitational field [12]. For these reasons, it would be interesting to analyse the behaviour of experimentally viable quantum states with background Schwarzschild or Kerr metrics as a way to unveil or predict some further, hopefully surprising, connections between classical gravity and the microscopic world. Apart from employing Maxwell’s covariant equations, it would also be interesting to analyse predictions of quantum field theory on a fixed background. The WKB approximation would be applied to obtain results in a general metric. Work along these lines is in progress.

**References**


